Nova Herculis Is Again Increasing in Brilliance

Famous Exploding Star, First Brilliant At Christmas Time, Reaches Eighth Magnitude After Waning To 13th

OVA Herculis, the newly discovered nova star whose peculiar astronomical antics have caused wide-spread interest within the past few months, is again increasing in brilliance, Leon Campbell of the Harvard College Observatory staff announced.

Rallying after it reached its faintest magnitude, 13, early in May, the star has now reached the eighth magnitude, has undergone marked spectrum changes and has definitely passed into what is usually termed the nebular stage, he

"It now looks as if the nova had passed through its stages of marked, and irregular, activity and that it will soon settle down to a point where it will slowly, but surely, decrease to the magnitude it was before its sudden outburst in December last," Mr. Campbell said. "The return to normal may require several years."

The star was discovered Dec. 13, 1934, by the English meteor observer, J. M. P. Prentice, and at that time was of the third magnitude, having flashed up from the fourteenth magnitude, probably, within a very few days, according to Mr. Campbell. (SNL, Dec. 22, 1934). From then until Dec. 22 the light of the nova steadily increased, with slight fluctuations, and on Dec. 22 the star reached the magnitude of 1.4 and was exceeded in brightness only by brilliant Vega.

Until April 1, the star was easily seen with the naked eye, varying between the second and fourth magnitudes, and then in two days it faded away at the rapid rate of about a magnitude and a half a day. After April 3 the decrease was more gradual, according to Mr. Campbell, until in early May it reached its faintest magnitude, 13, where it was seen only with the greatest difficulty with powerful telescopes.

The star is now clearly visible in small telescopes at any time throughout the night, passing through the zenith about midnight.

"Practically all the changes in brightness have been accompanied by changes in spectrum," Mr. Campbell said. "During the early stages, the nova was of a distinctly bluish color, indicating that it was then an extremely hot star.

When first photographed, the nova showed a spectrum containing strong absorption lines of hydrogen, with the lines of the metals bordered by intense emission lines. As the nova became brighter the emission lines faded away and the star had a spectrum closely resembling that of the supergiant star, Gamma Cygni.

"After maximum light was reached, the emission lines flashed forth with customary nova brilliance, this indicating that the nebulous shell surrounding the star had become partially transparent. Later, the well-known green auroral lines appeared as have been found in many other novae.

"Since the comeback in the star's light, marked spectrum changes have occurred. The nova has definitely passed into what is usually termed the nebular stage. The spectrum is outstanding in many ways, even among novae, and therefore difficult of interpretation at

'How bright will nova Herculis get? Judging by its prototype, Nova Aurigae 1891, probably it will not attain naked eye visibility but it will be visible in moderate sized telescopes."

At any rate, most of its activity is probably over, Mr. Campbell said, and it will soon settle down to a point where it will slowly decrease to the magnitude it was before its sudden outburst last December.

Science News Letter, July 13, 1935

Strange Behavior of Nova **Baffles Scientists**

'HAT Nova Herculis has science baffled in trying to explain its behavior, is the outstanding conclusion of a symposium, held in connection with the meeting of the Pacific division of the American Association for the Advancement of Science.

But while now puzzled, the astrono-

mers are going to do something about it. New instruments are being planned which will carefully scan the sky for the appearance of similar nova outbursts in the future and find them quicker. With the anticipated increase in observations of novae, the physical laws which govern their apparent erratic behavior may become known.

The astronomers who have been observing Nova Herculis for the last seven months have watched it wax and wane in brightness. A check on the "life histories" of former nova stars revealed that it has behaved like none of them.

All the astronomers can agree on now is that there is no standard behavior for an exploding star. Each is unique and a law unto itself. They all flash up quickly and then die down more slowly, but not in uniform fashion. Their antics cannot be predicted with knowledge now at hand.

It is the lack of knowledge which has led to plans for "nova finders" which, while still in the "paper" stage, may soon bring science out of its ignorance in the study of exploding stars.

Science News Letter, July 13, 1935

Extra High Voltage Not Needed To Cause Lightning

THE ENORMOUSLY high electric voltage which science has always associated with lightning bolts appears to be unnecessary.

Lightning progresses through a series of steplike jumps, and each jump requires only a part of the millions of volts potential necessary to jump from a high cloud to the ground.

Dr. B. F. J. Schonland, D. J. Malan and H. Collins astounded the learned Royal Society of London with the experimental proof so upsetting to science's previous conceptions of lightning's be-

Eighty-year-old Sir Charles V. Boys, inventor of the special high-speed camera which Dr. Schonland and his colleagues used in the new work, said the discovery was "an amazing phenomenon which could never have been predicted." (See SNL., Mar. 17, 1934, for report of former experiments.)

It is found that when a multiple lightning flash occurs there comes first a slow-moving "stepped leader," which precedes the first flash. Compared with subsequent happenings in the lightning stroke, this preliminary "leader" is slow, lasting as long as a fiftieth of a second. It is possible to overlook this action when the Boys camera runs at its fastest speeds.