

acids into protein molecules, contain special RNA. Researchers have suggested that before protein synthesis, nucleotides of ribosomal RNA must bind temporarily to nucleotides of messenger RNA, the intermediary between DNA and protein formation. When Sanger and colleagues examined the nucleotide sequences before the start signal of each phiX174 protein, they found that messenger RNA would have a stretch of at least four nucleotides that could bind with the nucleotides at one end of the ribosomal RNA. The sites where ribosomes bind have also been isolated from pieces of messenger RNA both in Sanger's laboratory by Nigel L. Brown and Michael Smith and by Rockefeller University biologists Jeffrey V. Ravetch, Peter Model and Hugh D. Robertson. Two articles also in the Feb. 24 NATURE describe the sequences of these sites and locate the sites on the overall phiX174 map. In all these ribosome binding sites there is a sequence of nucleotides that should be able to form stable base pairs with the end of the ribosomal RNA. The number of these possible base pairs, however, does not reflect how frequently ribosomes bind to each site and make the appropriate protein, so other factors must be involved in control of protein synthe-

sis.

The sequence has certainly not answered all questions. For example, the researchers learned little about what is special about the position on the DNA molecule, the replication origin, where copies of the viral DNA always begin. Sanger found that this origin falls within the gene for protein A, but that area showed no obvious symmetry or special shape.

Furthermore, the analysis opened the possibility of there being even more phiX174 proteins. "With the presence of two pairs of overlapping genes, the genome has more coding capacity than had been originally supposed on the assumption that each gene was physically separate," Sanger says. He points out that there are sites in the A, F, G and H genes where sequences for other proteins could possibly start.

Nobody knows whether the mechanisms uncovered by solving the phiX174 sequence will also appear in organisms that are less limited in total DNA. But molecular biologists now have new phenomena to watch for and plenty to think about. Everything the virus phiX174 is or does should be explicable beginning with that DNA sequence. □

One man's Mars: No Martians

For months, the Viking project's biologists have been struggling to find a theory that would explain the seemingly conflicting results of their experiments on the Martian surface. Now the first such theory has been offered, and, says Vance Oyama of the National Aeronautics and Space Administration's Ames Research Center, it shows "no need to invoke biological processes."

It is far from universally accepted, and it requires invoking several specialized chemical processes such as catalysis and polymerization. But, says Oyama, it fits all the facts, and elegantly so. Furthermore, if Oyama's theory is correct, it bears not only on the biology-instrument results, but also on the nature of the magnetic material that clings to Viking's test magnets, as well as on atmosphere-surface interactions and even on the early evolution of the planet.

The theory begins with the simplest of photochemical effects in the atmosphere—the dissociation of the dominant carbon dioxide by solar ultraviolet radiation into "activated" carbon monoxide and single atoms of oxygen. With more UV, some of the CO is further dissociated into carbon and oxygen. The carbon then combines with CO to form carbene (C₂O), which in turn combines with still more CO to form the first key element in Oyama's proposal, carbon suboxide, or C₃O₂. The carbon suboxide molecules, Oyama says, can then unite into a carbon suboxide polymer, (C₃O₂)_N. All of these reactions,

he adds, are known from laboratory studies, and the resulting polymer has a reddish cast very much like that of the Martian surface.

The next step is to plug this result into the data from the three different kinds of biology experiments aboard the landers. Oyama's explanation begins with the pyrolytic-release instrument—the one of the three that has suggested reduction processes.

The PR instrument works by exposing a soil sample to an atmosphere that includes CO and CO₂ whose carbon atoms are radioactive carbon 14. After an incubation period, the sample is incinerated, or pyrolyzed, and the resultant gases are monitored for radioactivity as an indication of how much of the CO and/or CO₂ was incorporated into the soil.

During the design of the Viking experiments, it had been suggested by some observers that the radioactive carbon isotopes might kill off the very microorganisms that they were trying to measure. Viking's biologists have long felt that the radioactivity levels are too low to do any damage, and Oyama sees no evidence of life to kill anyway, but the isotopes do play an important role in his explanation of the PR results.

The decay of the carbon 14 (into nitrogen 14) releases a beta particle, he points out, and the energy of this decay product (0.156 MeV, Oyama says), multiplied by the 22 microcuries of radioactivity in each PR "cycle," is more than sufficient to

The closest look yet at Phobos

Viking orbiter 1 has provided the closest look ever taken at the Martian moon Phobos (see cover), revealing a wealth of detail and enabling the first accurate calculations of the tiny satellite's density. A very preliminary look, says Joseph Veverka of Cornell University, suggests about 2 grams per cubic centimeter, consistent with earlier speculation that Phobos resembles carbonaceous chondritic material. Veverka and others feel that this is evidence that Phobos (and Deimos) may be captured objects originally formed far out in the asteroid belt between Mars and Jupiter. The photos have shown many small, bowl-shaped craters which, combined with the flatter floors of larger craters, should yield an estimate of the depth of the planetoid's regolith, or fragmented surface, with substantial implications for its evolution and history. One of Phobos's major mysteries—a series of strange, parallel striations crossing part of the surface—shows up clearly in the closest photos as grooves, about 100 to 200 meters wide and tens of kilometers long, rather than as "chains" of small craters. (Such chains have also been seen, suggesting secondary ejecta from larger craters.) The grooves do not appear on every side. This implies that they are symptoms neither of Mars-caused tidal distortions nor of strata seen edge-on. They could be related to the formation of Phobos's largest crater, Stickney. This will seem even more likely if the other moon, Deimos, has none. □

fracture carbon-carbon, carbon-hydrogen or carbon-oxygen bonds. This breakage has the effect of "activating" the red carbon suboxide polymer, making it "receptive" to incorporating the provided carbon monoxide. When the polymer is heated to 625°C, it yields about 4 percent of itself as the original, but now labeled, monomer form, which sticks to the experiment's organic-vapor trap and appears, upon further heating, as the critical "second peak" in the experiment's data.

If water vapor is present in the PR experiment when the sample is exposed to the labeled atmosphere, says Oyama, this second peak is lower. In Oyama's own "gas-exchange" experiment, which exposes a sample to a nonlabeled nutrient solution and simply monitors changes in the composition of the test-cell atmosphere, the prominent release of oxygen is also reduced. But the reason, Oyama says, is very different.

In the Martian atmosphere, the same photodissociative reactions which lead to the formation of carbon suboxide also lead, by another path, to activated oxygen